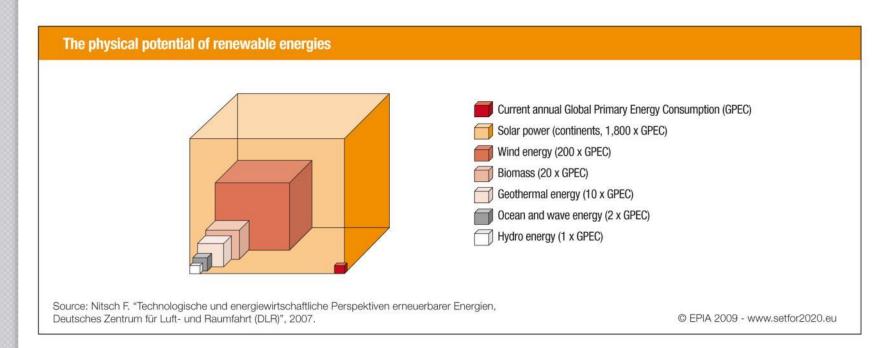




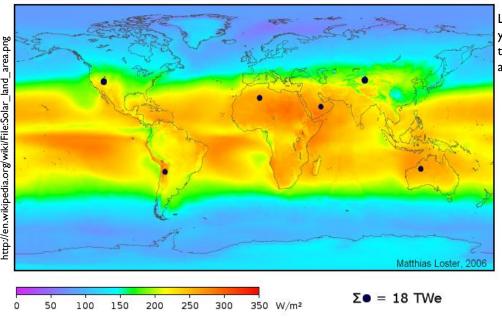
- Radiation from the Sun
- Atmospheric effects
- Insolation maps
- Tracking the Sun
- PV in urban environment

- Solar resource is immense
 - Human energy use: 4.0×10¹⁴ kWh/year
 - Solar resource on Earth's surface: 5.5×10¹⁷ kWh/year





- Solar resource is immense
 - Human energy use: 4.0×10¹⁴ kWh/year
 - Solar resource on Earth's surface: 5.5×10¹⁷ kWh/year

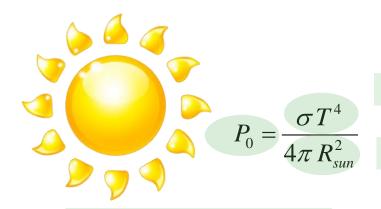


Local solar irradiance averaged over three years from 1991 to 1993 (24 hours a day) taking into account the cloud coverage available from weather satellites

Solar power systems covering the areas defined by the dark disks could provide more than the world's total primary energy demand (assuming a conversion efficiency of 8%).



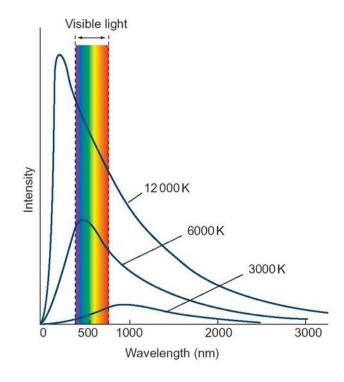
- Solar resource is immense
 - Human energy use: 4.0×10¹⁴ kWh/year
 - Solar resource on Earth's surface: 5.5×10¹⁷ kWh/year
- Solar resource is global and democratic
- Solar resource is relatively constant but depends on
 - atmospheric effects, including absorption and scattering
 - local variations in the atmosphere, such as water vapour, clouds, and pollution
 - latitude of the location
 - the season of the year and the time of day

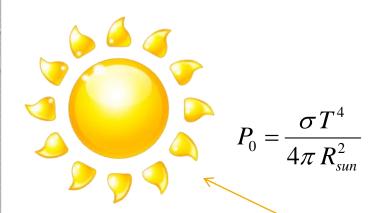


Power radiated per unit area $5.96 \times 10^7 W/m^2$

Total radiative power (Stefan Boltzman) T=5762K

Surface area of sun





Ratio of surface areas of the 2 spheres

Solar constant average energy flux incident at the Earth's orbit: 1366 W/m²

$$S = \frac{4\pi R_{sun}^2}{4\pi D^2} P_0$$



Energy incident on Earth

Total area of Earth



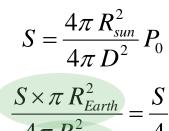
$$P_0 = \frac{\sigma T^4}{4\pi R_{sun}^2}$$

6.96x105 km R_{sun}

 $\mathsf{D}_{\mathsf{avg}}$ $1.5 \times 10^8 \, \text{km}$

 R_{Earth} $6.35 \times 10^{3} \text{ km}$

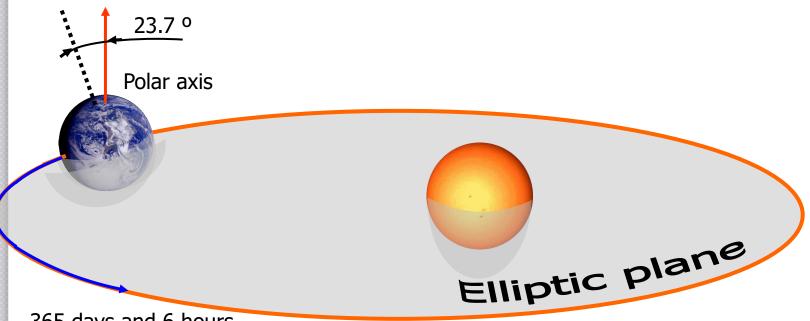
 $\frac{S \times \pi R_{Earth}^2}{4\pi R_{Earth}^2} = \frac{S}{4}$





Average energy incident per unit area of surface of Earth: 342 W/m²

Earth-Sun motion

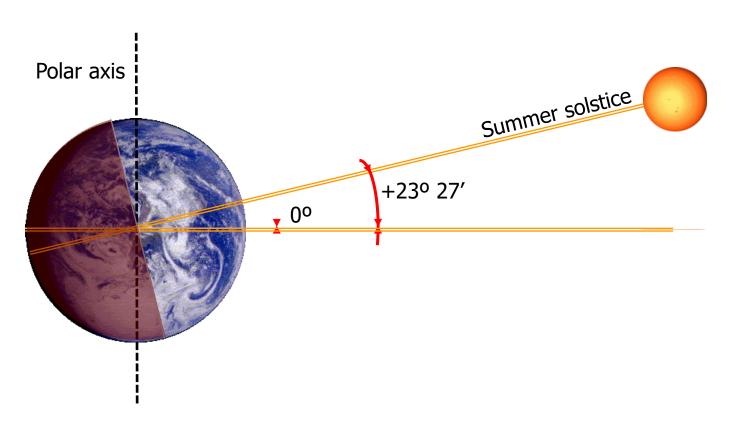


365 days and 6 hours

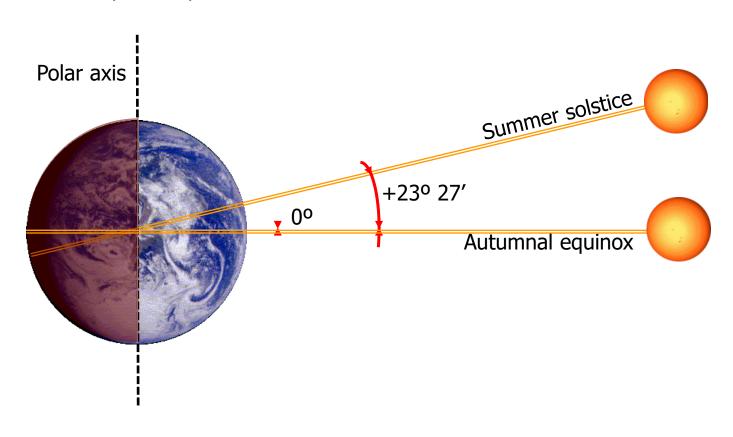
$$\frac{H}{S} = 1 + 0.033 \cos\left(\frac{360(n-2)}{365}\right)$$

H(W/m2) is radiant power density outside the atmosphere; S is solar constant; n is day of the year

- Earth-Sun motion
 - Solar declination: angle between line joining centres of Earth and Sun and the equatorial plane

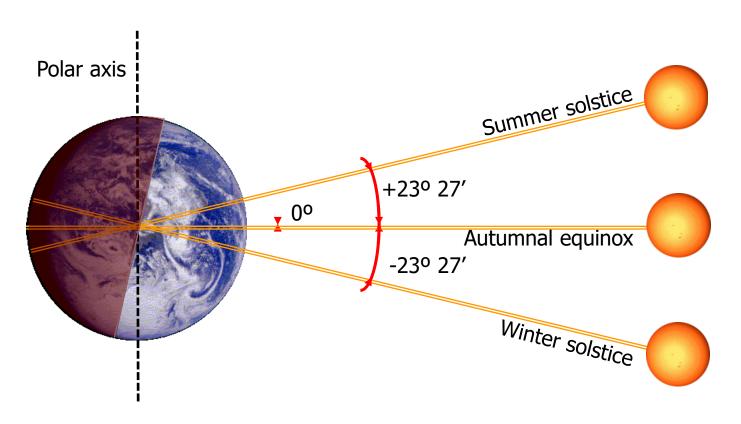


- Earth-Sun motion
 - Solar declination: angle between line joining centres of Earth and Sun and the equatorial plane



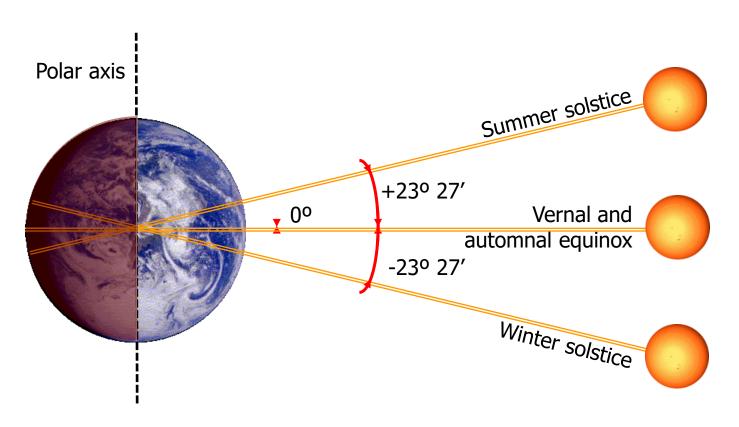
• Earth-Sun motion

 Solar declination: angle between line joining centres of Earth and Sun and the equatorial plane



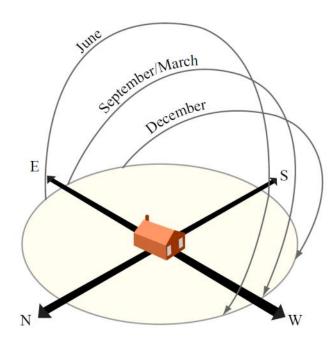
• Earth-Sun motion

 Solar declination: angle between line joining centres of Earth and Sun and the equatorial plane



Earth-Sun motion

 Solar declination: angle between line joining centres of Earth and Sun and the equatorial plane

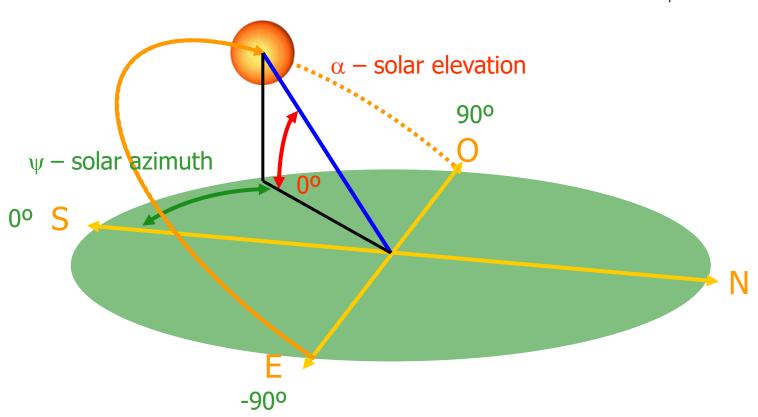


Building orientation with the long axis facing south

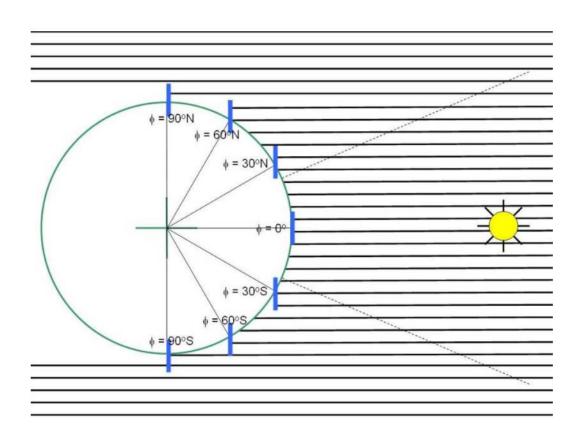
$$\delta = \pi \frac{23.45}{180} \sin \left(2\pi \frac{284 + n}{365} \right)$$

• Earth-Sun motion

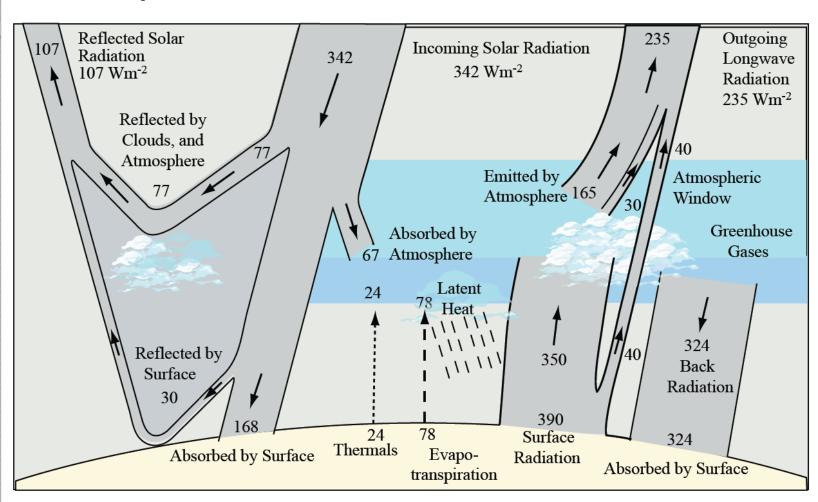
$$\sin \alpha = \sin \delta \sin \phi + \cos \delta \cos \phi$$
$$\cos \psi = \frac{\sin \alpha \sin \phi - \sin \delta}{\cos \alpha \cos \phi}$$



- Optimum orientation: facing south (north in the southern hemisphere)
- Optimum inclination: local latitude but not quite



Atmospheric effects

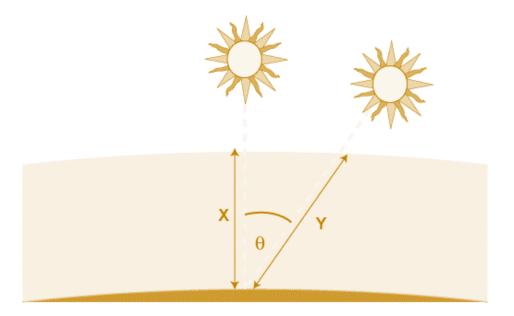




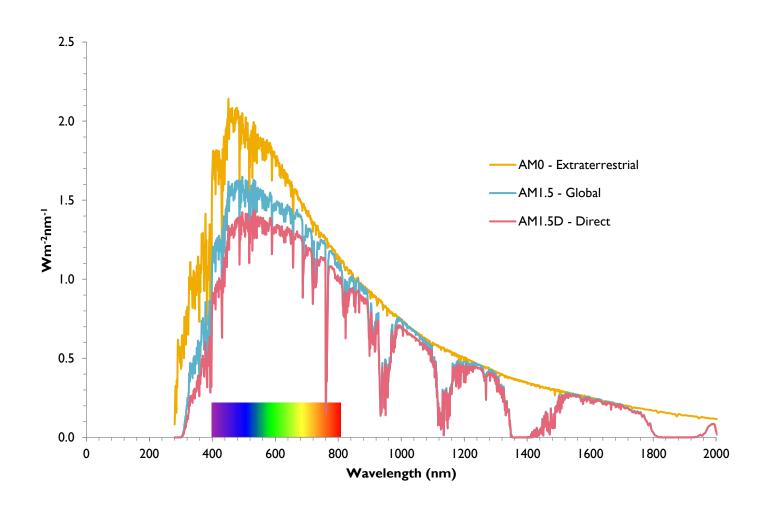
Atmospheric effects on solar radiation at the Earth's surface:

- a reduction in the power of the solar radiation due to absorption,
 scattering and reflection in the atmosphere;
- a change in the spectral content of the solar radiation due to greater absorption or scattering of some wavelengths;
- the introduction of a diffuse or indirect component into the solar radiation; and
- local variations in the atmosphere (such as water vapour, clouds and pollution) which have additional effects on the incident power, spectrum and directionality.

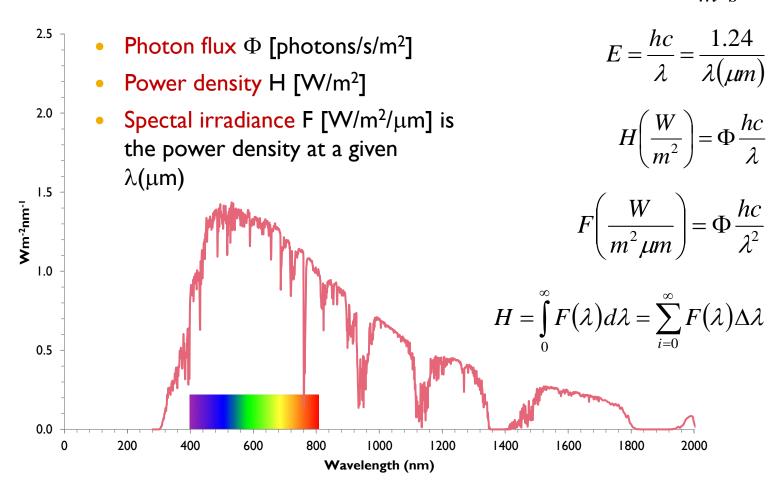
 Air Mass is a measure of the reduction in the power of light as it passes through the atmosphere and is absorbed by air and dust

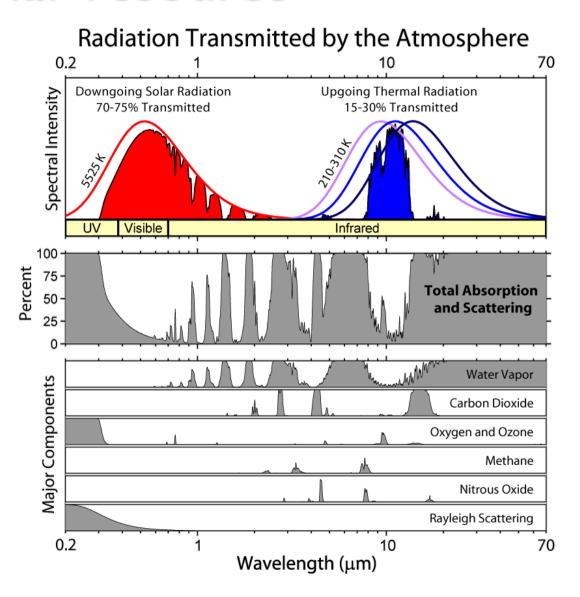


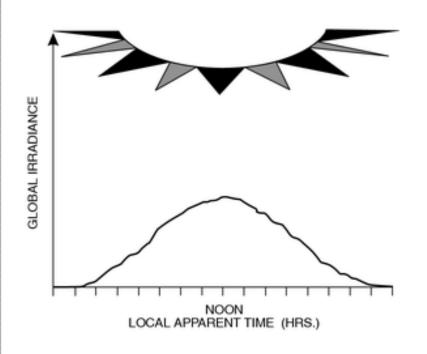
$$AM = \frac{1}{\cos \theta}$$

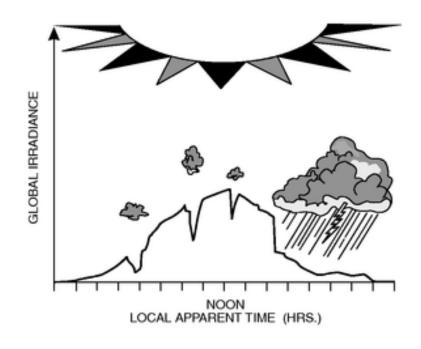


$$\Phi = \frac{photons}{m^2s}$$





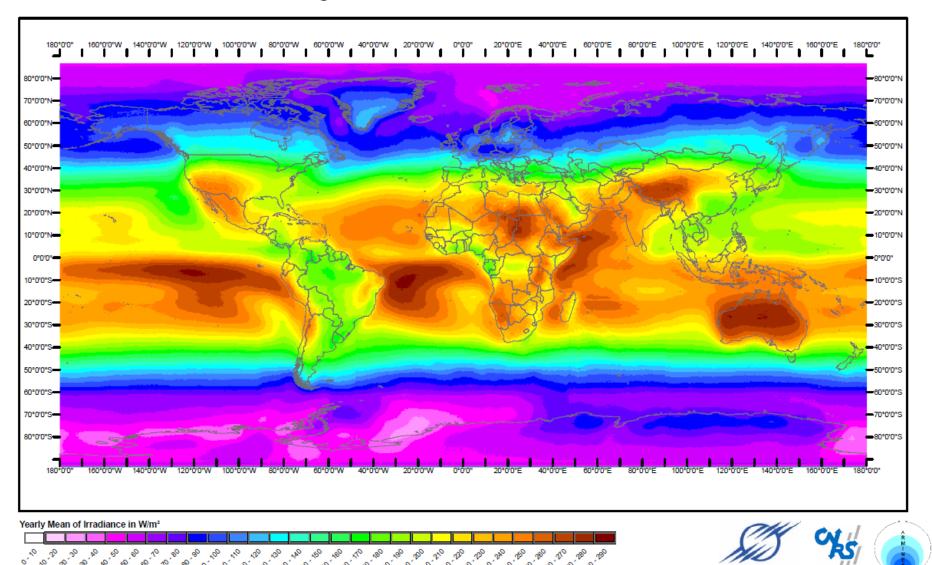






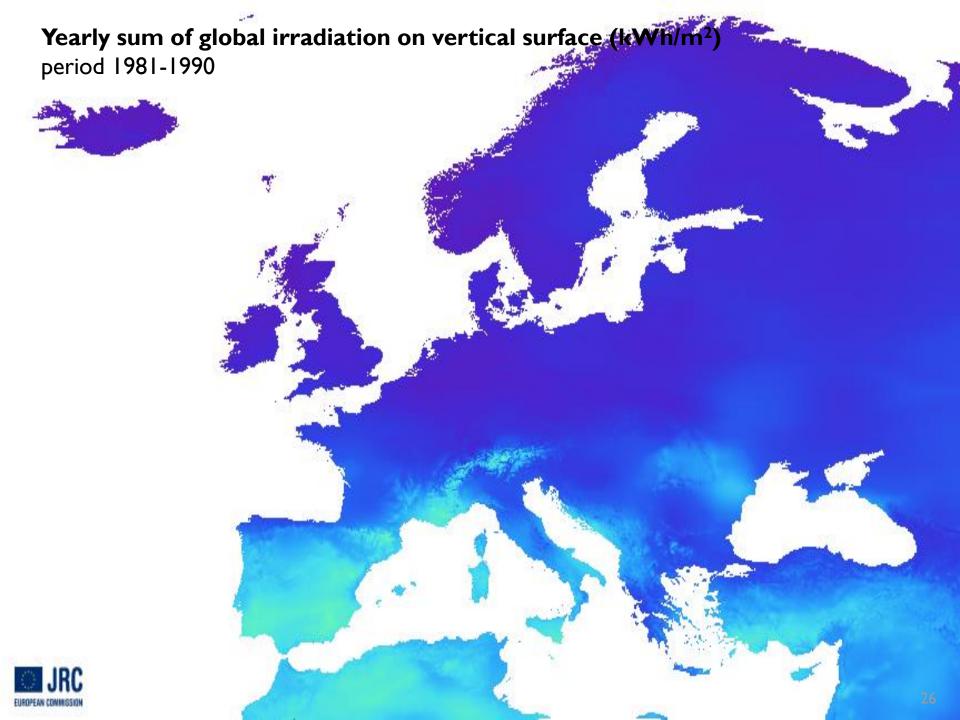
- Insolation: Incoming Solar Radiation
- Typical units: kWh/m²/day
- Affected by latitude, local weather patterns,...
- Šúri M., Huld T.A., Dunlop E.D. Ossenbrink H.A., 2007. Potential of solar electricity generation in the European Union member states and candidate countries. Solar Energy, 81, 1295–1305, http://re.jrc.ec.europa.eu/pvgis/.

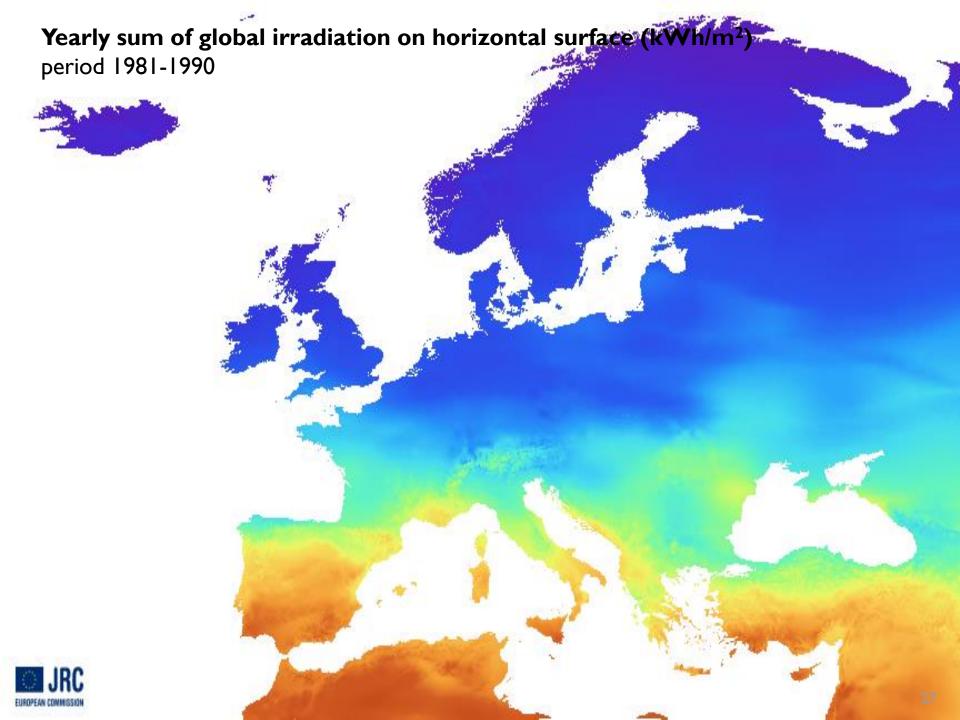
Averaged Solar Radiation 1990-2004

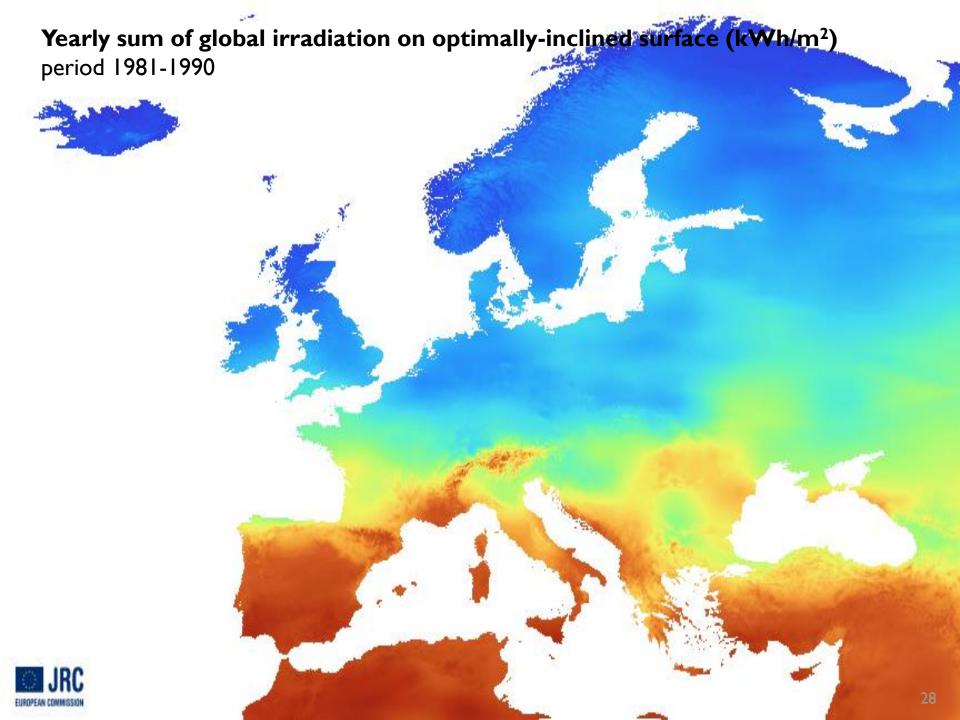


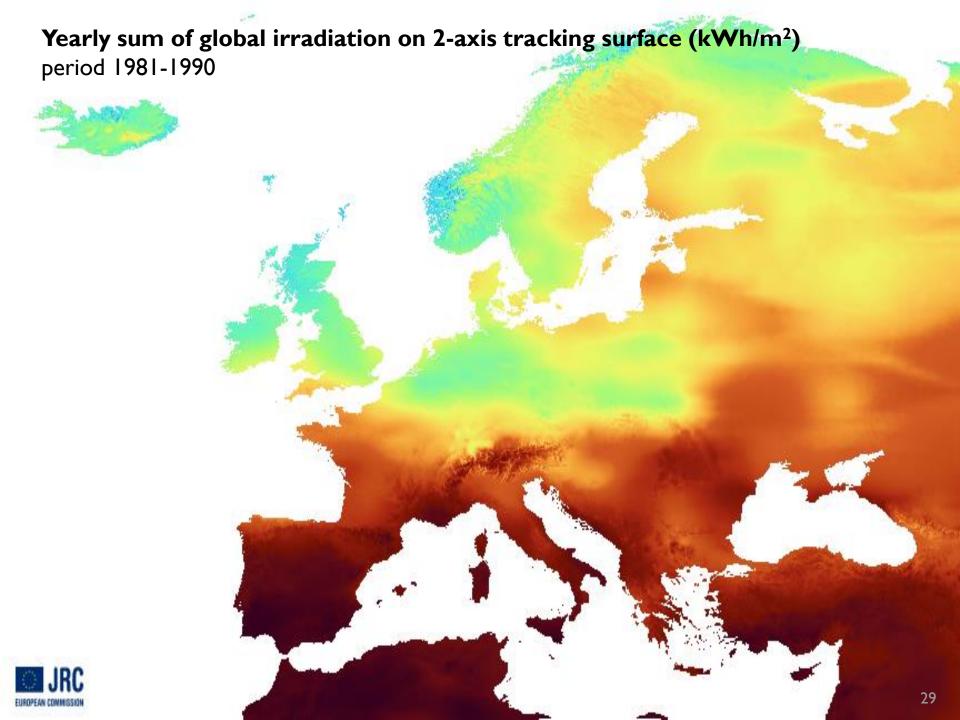
Realized by Michel Albuisson, Mireille Lefèvre, Lucien Wald.
Edited and produced by Thierry Ranchin. Date of production: 23 November 2006.
Centre for Energy and Processes, Ecole des Mines de Paris / Armines / CNRS.
Copyright: Ecole des Mines de Paris / Armines 2006. All rights reserved.

MINES PARIS





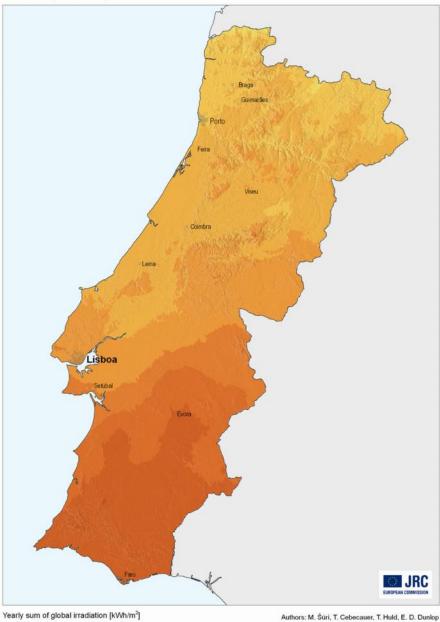


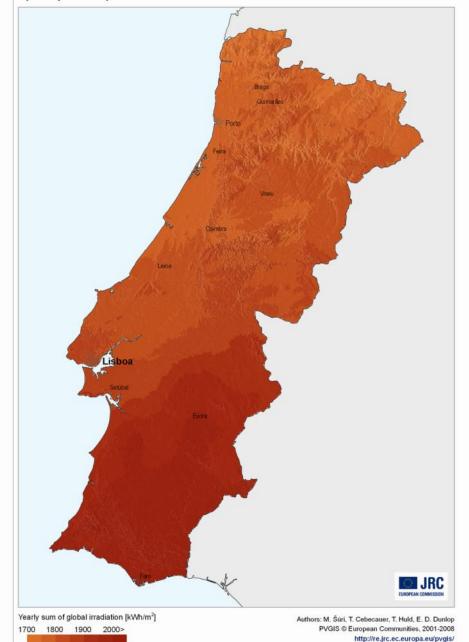


1275 1350 1425 1500>

Yearly electricity generated by 1kW_{peak} system with performance ratio 0.75 [kWh/kW_{peak}]

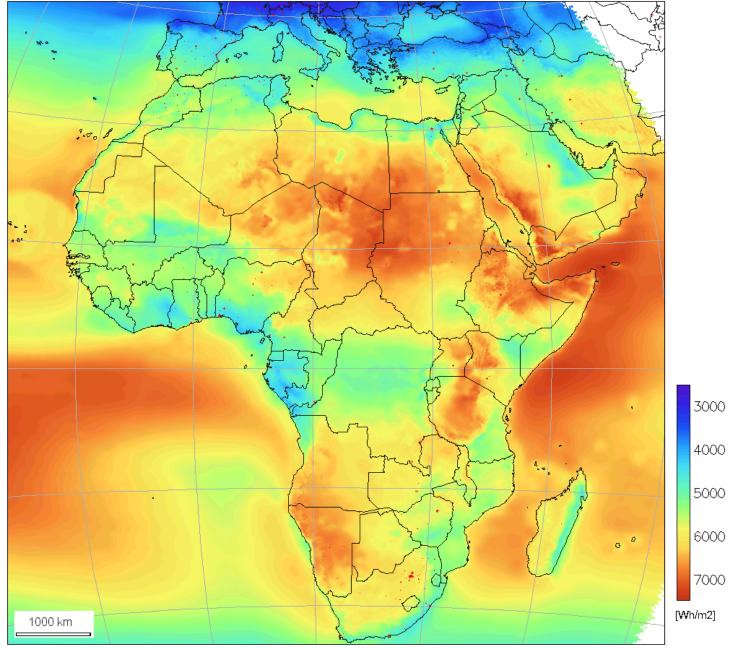
25











1000 1200 1400 1600 1800 2000

Figure 1: Long-term average of yearly sums of global horizontal irradiation (kWh/m², time series representing years 1985, 1987, and 1989-2004)

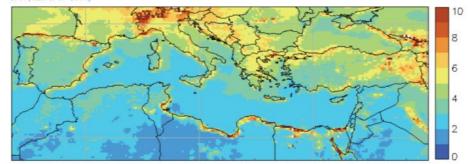


Figure 2: Standard deviation (in %) of yearly sums of global horizontal irradiation (years 1985, 1987, and 1989-2004)

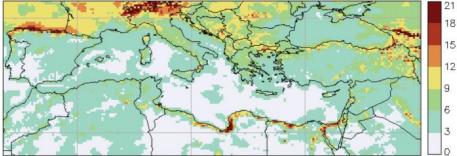


Figure 6: Relative difference of the highest yearly sum of global horizontal irradiation in relation to the long-term average (in %)

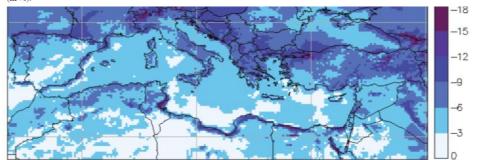


Figure 5: Relative difference of the lowest yearly sum of global horizontal irradiation in relation to the long-term average

Solar resource

- Coastal areas and higher mountains face wider variations (up to 10%)
- Winter is much more variable (up to x6) than summer months

Šúri M., Huld T., Dunlop E.D., Albuisson M., Lefèvre M., Wald L., 2007. Uncertainties in photovoltaic electricity yield prediction from fluctuation of solar radiation. Proceedings of the 22nd European Photovoltaic Solar Energy Conference, Milano, Italy 3-7.9.2007

Solar tracking

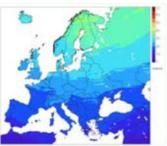
Compared to PV with modules fixed at optimum angle:

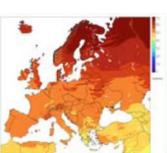
Changing inclination twice a year contributes only marginally

Fixed mounting - two (seasonal) optimum angles









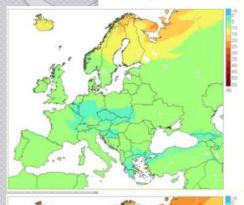
Optimum angle of modules in summer [°]

+ 1.5 to 4.5 %

Optimum angle of modules in winter [°]

Solar tracking

One-axis tracking



Horizontal axis pointing East-West



Vertical axis + modules mounted at optimum angle





Optimum angle of modules [°]

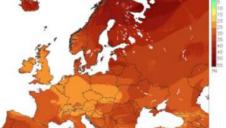
+ 11 to 55 %





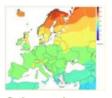
pointing North-South





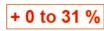
Axis inclined at an optimum angle towards South





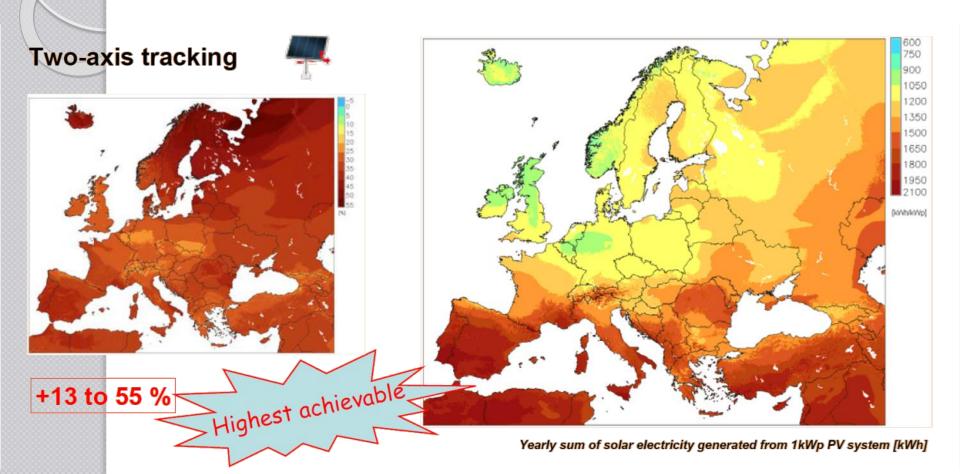


Optimum angle of the axis [°]



Horizontal axis

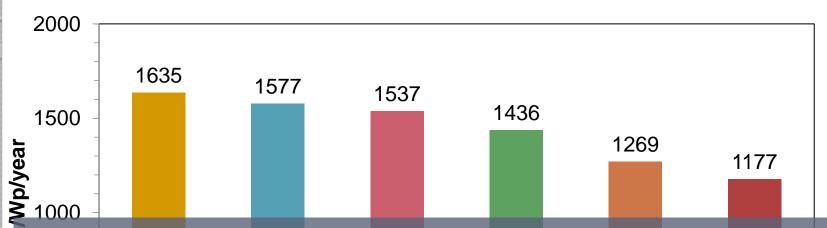
Solar tracking





Compared to PV with modules fixed at optimum angle:

- Changing inclination twice a year contributes only marginally (2-4%)
- I-axis tracking PV with vertical or South-inclined axis generates only I-4% less than 2-axis tracking system
- I-axis tracking PV with horizontal axis-oriented E-W typically performs only slightly better than fixed mounting systems



- An inclined axis tracking system orientated in the north-south direction is able to closely follow the Sun throughout the year, and have a relatively high acceptance for CPV applications.
 - The vertical tracking, with optimum inclination, features a similar performance that an inclined axis tracking system orientated in the north-south direction.







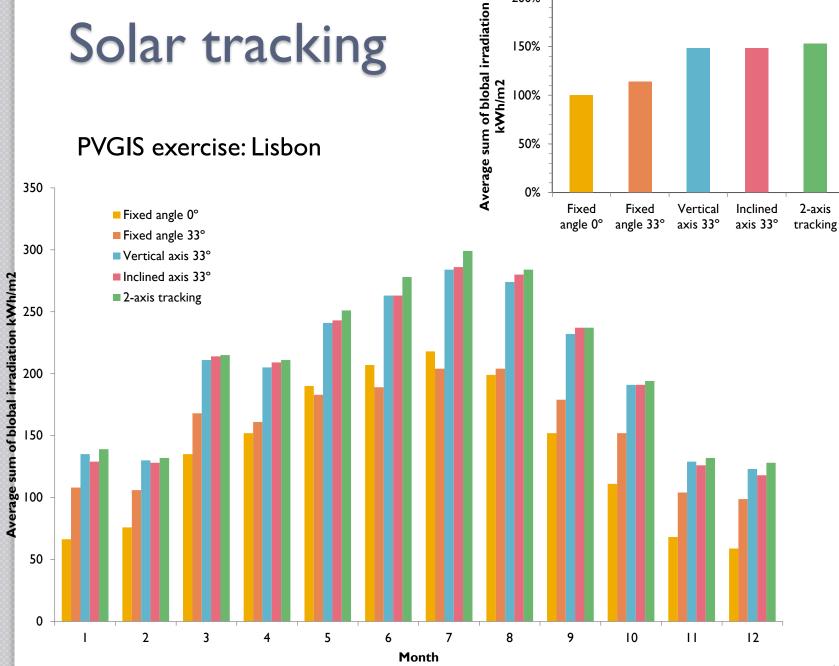




▲ Installing a screw foundation for a dual-axis tracking system (Deger-Traker) at Am Peterswald PV park in Germany: Bftec GmbH pre-drills the ground at the site and then the foundation is installed. The foundation's unusual form, which consists of a smooth side wall with a coil at the bottom (if necessary, two coils), guarantees a firm grip. The foundation can rotate in the ground without displacing or loosening the surrounding earth.



PVGIS exercise: Lisbon



200%

150%

00%

Shadowing effect

Ground cover ratio = PV area / total area

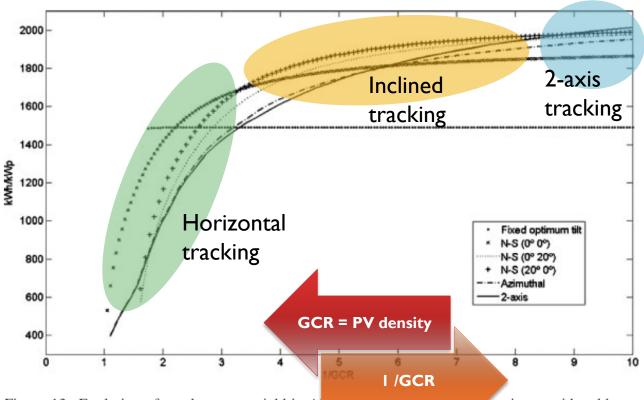
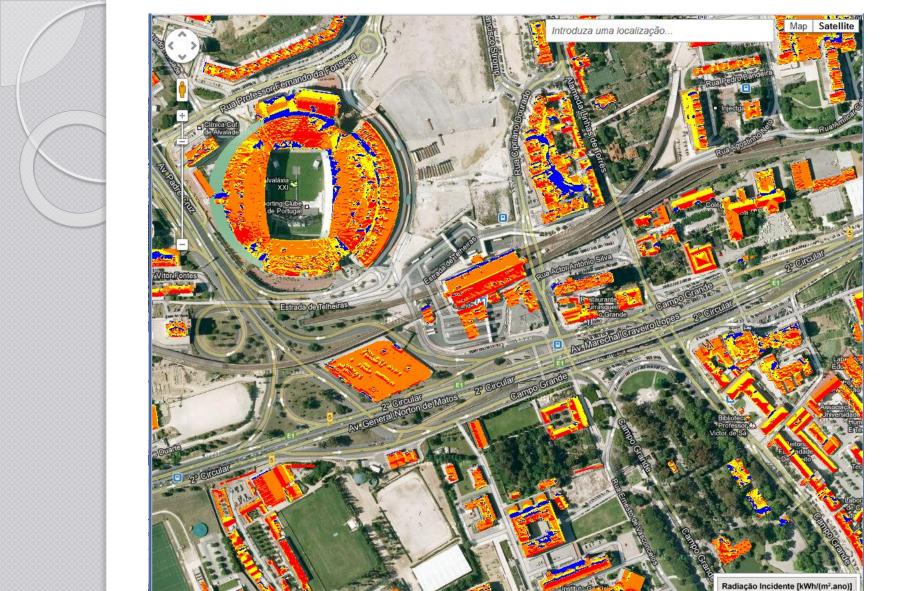


Figure 13. Evolution of yearly energy yield in Almeratics are attegies considered here, for the pessimistic shading case and assuming a constant dir ness degree of 3%

PV potential in the urban landscape is harder to estimate

- ☐ Geographical solar potential locations where this energy can be captured
- ☐ **Technical** solar potential technical characteristics of the rooftop/equipment used
- Economic solar potential only viable systems

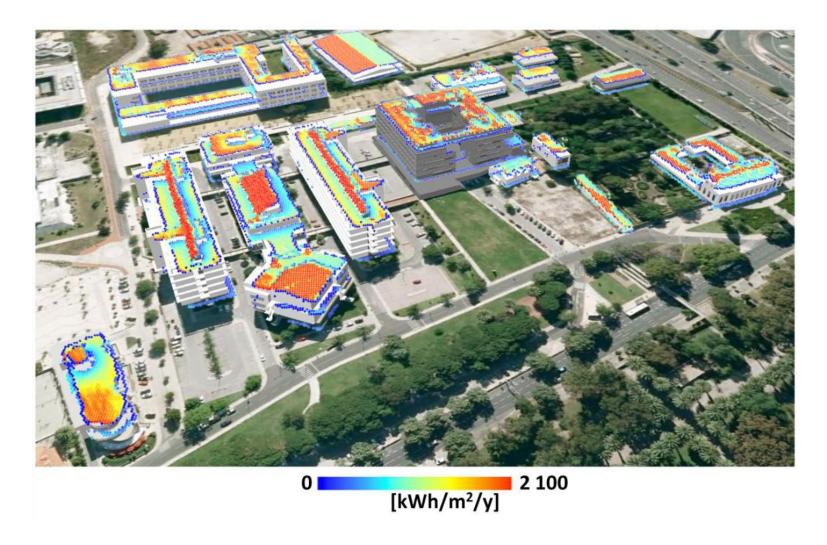


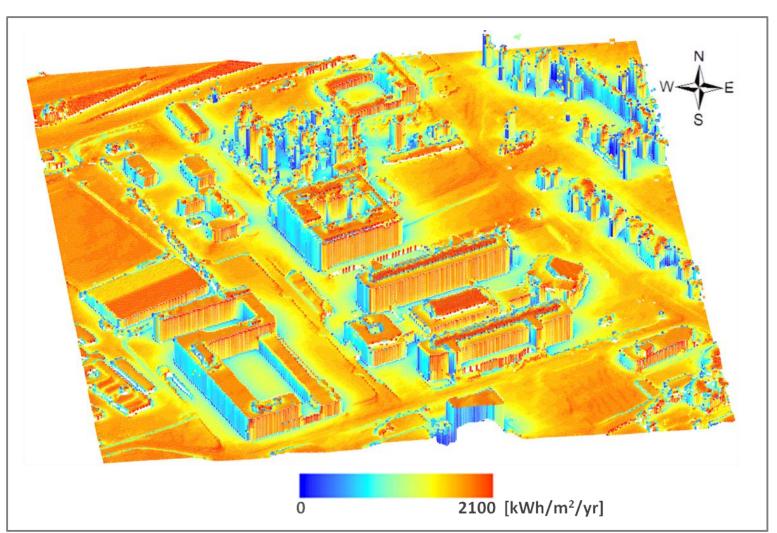


CLASSE III]1400-1600]

LISBOA e·nova







Redweik et al, Solar Energy 97 (2013) 332–341



